

## Normal Modes of Self-gravitating Fluids in Perturbed Configurations

### II. Perturbational-variational Expansion of the $g$ - and $p$ -modes of a Non-adiabatic Fluid about the Adiabatic Limit\*

Y. Sobouti<sup>1</sup> and J. N. Silverman<sup>2</sup>

<sup>1</sup> Department of Physics and Biruni Observatory, Pahlavi University, Shiraz, Iran

<sup>2</sup> Department of Physics and Molecular Physics and Quantum Chemistry Group, Pahlavi University, Shiraz, Iran

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**Summary.** The  $g$ -eigenfrequencies of an adiabatic fluid are identically zero; their growth rate with departure of the fluid from the adiabatic limit and the subsequent motions, however, are obtainable from an eigenvalue equation. The  $p$ -modes of the adiabatic fluid are also solutions of another eigenvalue equation. These two eigenvalue problems are derived and solved. In a non-adiabatic fluid, taking the departure of the fluid from adiabaticity as a perturbation parameter and using a perturbational-variational Rayleigh-Ritz technique, the  $g$ - and  $p$ -modes are expanded about those of the adiabatic limit mentioned above. The expressions for the zeroth- and the first-order  $g$ - and  $p$ -eigenvalues and eigenvectors, and the second-order  $g$ -eigenvalues are analyzed and computed. With regard to convection, the information on  $g$ -modes should be of particular interest: In a *slightly* superadiabatic fluid, the first-order  $g$ -eigenvalues and their corresponding eigenvectors give the time rate of growth of the convective instabilities and the patterns of convective motions, respectively.

**Key words:** generalized perturbed eigenvalue equation — perturbational-variational Rayleigh-Ritz — self-gravitating fluids: normal modes, convection, pulsation — stellar interiors

### I. Introduction

An immediate inference from the Schwarzschild stability criterion is that adiabatic fluids in gravitational fields are in neutral convective equilibrium. That, however, neutral convective motions could exist in the fluid was mathematically confirmed by Lebovitz (1965a, b and 1966). Sobouti (1977a, b), identifying these motions of the adiabatic fluid with its neutral  $g$ -modes, proposed a definition and developed a mathematical expression for the  $g$ -modes of self-gravitating fluids. The definition and

the appropriate representation for the  $p$ -modes then followed from the requirement that the  $p$ -modes should be orthogonal to the  $g$ -modes. In the latter formalism, adiabatic fluids prove to have a simpler structure of normal modes than the non-adiabatic media. In fact, an analytical separation of the space of normal modes of adiabatic fluids into a  $g$ - and a  $p$ -subspace becomes possible.

In the present analysis, we expand the normal modes of a non-adiabatic fluid about those of an adiabatic structure. A measure of departure from adiabaticity is taken as the perturbation parameter and the procedure of Silverman and Sobouti (1977, henceforth referred to as Paper I) is employed to obtain a series expansion of the eigenfrequencies and eigendisplacements. The analysis is carried out within the framework of Sobouti's definition of the  $g$ - and  $p$ -modes and the identification of the  $g$ -modes with convective oscillations. The latter work (Sobouti, 1977a) will be referred to as Paper III. Frequent references to equations of Papers I and III will be indicated by the roman numerals I or III, respectively, before the equation number in question; for example, Equation (I.7) will mean Equation (7) of Paper I.

In Section II, the equation of motion and its matrix representation are introduced. In Section III, the perturbation parameter and the expansion of various operators and their equivalent matrices are discussed. In Section IV, the procedure of Paper I is further extended to incorporate the bispectral  $g$ - and  $p$ -character of the normal modes into the formalism. In Section V, the numerical procedure is analyzed, and in Section VI, conclusions and numerical results are discussed.

### II. Equation of Motion

Adiabatic Lagrangian displacements,  $\xi(r) \exp(i\epsilon^{1/2}t)$  of a self-gravitating fluid satisfy the following generalized eigenvalue equation,

$$\mathcal{W} \xi^s = \epsilon^s \rho \xi^s; \quad s = 1, 2, \dots, \quad (1a)$$

Send offprint requests to: J. N. Silverman

\* Contribution No. 2, Biruni Observatory

where

$$\mathcal{W}\xi = \nabla(\delta p) - \frac{1}{\rho}\delta\rho\nabla p - \rho\nabla(\delta\Omega), \quad (1b)$$

$$\delta p = -\gamma p\nabla\cdot\xi - \xi\cdot\nabla p, \quad (2a)$$

$$\delta\rho = -\rho\nabla\cdot\xi - \xi\cdot\nabla\rho, \quad (2b)$$

$$\nabla^2(\delta\Omega) = -4\pi G\delta\rho, \quad (2c)$$

and  $\rho$ ,  $p$  and  $\Omega$  denote the density, the pressure, and the gravitational potential, respectively.

In the Rayleigh-Ritz approximation and in the notation of Paper I, the eigendisplacements,  $\xi^s$ , will be expanded in terms of a finite set of basis vectors,  $\{\zeta^t\}$ ;  $t=1, 2, \dots, n$ . Thus,

$$\xi^s = \sum_t^n \zeta^t Z^{ts}. \quad (3)$$

Equation (1a) will be written in its equivalent but approximate matrix form

$$WZ = SZE, \quad (4a)$$

with the corresponding orthonormality condition [cf. Eq. (I.7b)],

$$Z^\dagger SZ = I, \quad (4b)$$

where the  $W$ -,  $S$ -,  $Z$ - and  $E$ -matrices are defined by Equations (I.5) and (I.6). We note that  $W$  and  $S$  are real symmetric and  $S$  is also positive definite.

Solutions of Equations (1a) or (4) fall into two distinct classes of  $g$ - and  $p$ -modes. It was shown in Paper III that there exist two sets of basis vectors,  $\{\zeta_g^u\}$  and  $\{\zeta_p^v\}$ , which together span the Hilbert space of  $\{\xi^s\}$ . In the case of an adiabatic fluid,  $\{\zeta_g^u\}$  spans the  $g$ -subspace and  $\{\zeta_p^v\}$  spans the  $p$ -subspace of the normal modes. The two sets, in the sense of Equations (I.3b) or (III.6b), are orthogonal to each other, and an analytical separation of the  $g$ - and  $p$ -modes is possible. To be specific,  $\{\zeta_g^u\}$  consists of those displacements of the adiabatic fluid which leave the pressure equilibrium undisturbed; it follows from Equation (III.3) that they satisfy

$$-\gamma p_0 \nabla\cdot\zeta_g^u - \zeta_g^u\cdot\nabla p_0 = 0, \quad (5a)$$

where the subscript zero indicates that the function in question pertains to the adiabatic fluid. A single scalar function, say  $\nabla\cdot\zeta_g^u$ , is sufficient to specify the vector field  $\zeta_g^u$ . The  $\{\zeta_p^v\}$  are those displacements of the fluid which are orthogonal to  $\{\zeta_g^u\}$ ; they satisfy Equation (III.7),

$$\zeta_p^v = \nabla\phi^v, \quad (5b)$$

where  $\phi^v$ , again a single arbitrary scalar, specifies the vector field  $\{\zeta_p^v\}$ . In what follows, the truncated basis vectors  $\{\zeta_g^u; u=1, 2, \dots, n_g\}$  and  $\{\zeta_p^v; v=1, 2, \dots, n_p\}$ , combined together, will be used to expand the eigendisplacements,  $\xi^s$ , of the non-adiabatic fluid, Equation (3). It should be noted that the  $g$ -vectors of Equation (5a) depend on the pressure distribution of the *adiabatic* fluid,

while those of Equation (III.3) depended on the structure of the *non-adiabatic* fluid; as a consequence, the use of the former results in a slightly different set of matrix elements (cf. Appendix A) than those of Paper III. On the other hand, this restriction of the  $g$ -basis vectors to the adiabatic fluid renders the matrix elements independent of the perturbation parameter, which will be taken to be a measure of departure from adiabaticity, and is thus of fundamental consequence to the perturbational-variational procedure of Section III.

We now carry out the expansion of Equation (3) in terms of the  $g$ - and  $p$ -basis vectors discussed above. From Equation (III.8), this yields

$$\xi_g^s = \sum_u^{n_g} \zeta_g^u Z_{gg}^{us} + \sum_v^{n_p} \zeta_p^v Z_{pg}^{vs}, \quad (6a)$$

$$\xi_p^t = \sum_u^{n_g} \zeta_g^u Z_{gp}^{ut} + \sum_v^{n_p} \zeta_p^v Z_{pp}^{vt}. \quad (6b)$$

This partitioning of the eigendisplacements,  $\{\xi^s\}$ , and of the basis-vectors,  $\{\zeta^t\}$ , into  $g$ - and  $p$ -vectors results in a corresponding block-partitioning of all the matrices. Thus,

$$\{\xi^s\} = \{\xi_g^u, \xi_p^v\}, \quad (7a)$$

$$\{\zeta^s\} = \{\zeta_g^u, \zeta_p^v\}; \quad u=1, \dots, n_g; \quad (7b)$$

$$v=1, \dots, n_p; \quad s=1, \dots, n; \quad n=n_g+n_p,$$

and [cf. Eqs. (III.9)–(III.12)]

$$A = \begin{bmatrix} A_{gg} & A_{gp} \\ A_{pg} & A_{pp} \end{bmatrix}; \quad A = W, S, Z, E. \quad (7c)$$

By Equations (I.5), (III.11a) and (III.12a), the block elements of  $W$  and  $S$  are as follows:

$$W_{ab}^{ts} = \langle \zeta_a^t | \mathcal{W} | \zeta_b^s \rangle; \quad a, b = g, p, \quad (8a)$$

$$S_{ab}^{ts} = \langle \zeta_a^t | \rho | \zeta_b^s \rangle; \quad a, b = g, p. \quad (8b)$$

Explicit expressions for various elements of  $W$  and  $S$  are developed in Appendix A.

### III. Perturbational-variational (PV) Expansion

The unperturbed reference structure will be taken to be the adiabatic fluid, and the quantities pertaining to it will be indicated by a subscript zero. For the reference fluid, one has

$$\frac{\partial \ln p_0}{\partial \ln \rho_0} = \gamma, \quad \text{the ratio of specific heats}; \quad (9a)$$

the corresponding derivatives for a non-adiabatic fluid will naturally be different. The perturbation parameter,  $a$ , [i.e., the  $\lambda$  of Paper I, cf. Eqs. (I.2)] will be chosen as follows:

$$a = \frac{\partial \ln p}{\partial \ln \rho} \bigg/ \frac{\partial \ln p_0}{\partial \ln \rho_0} - 1. \quad (9b)$$

For superadiabatic, adiabatic, and subadiabatic fluids,  $a$  assumes positive, zero, and negative values, respectively; the parameter  $a$ , a measure of departure of the fluid from the adiabatic limit, is a Schwarzschild discriminant, and determines whether or not convection takes place. Furthermore, it was shown in Paper III that in the neighborhood of the adiabatic limit, the  $g$ -frequencies are proportional to  $a^{1/2}$ , thus becoming stable or unstable depending on whether  $a$  is positive or negative, respectively. For a polytrope of index  $n$ ,  $a$  has a constant value,

$$a = \frac{1}{\gamma} \left( 1 + \frac{1}{n} \right) - 1. \quad (9c)$$

Hereafter, the discussion will be confined to polytropes.

Assume that the structural parameters of a non-adiabatic fluid have the following convergent series expansions about the adiabatic limit:

$$\sigma = \sum_{j=0} \sigma_j a^j; \quad \sigma = p, \varrho, \nabla p, \nabla \varrho, \Omega, \quad (10a)$$

where

$$\sigma_j = \frac{1}{j!} \left( \frac{d^j \sigma}{da^j} \right)_{a=0}. \quad (10b)$$

In general, determination of  $\varrho_j$ ,  $p_j$ , etc., will require a study of the equilibrium structure of the fluid. For the present, assume the expansions of Equation (10a) are known; insert the latter in Equations (1b) and (8) to obtain corresponding expansions for the  $W$ - and  $S$ -matrices. Thus,

$$A = \sum_{j=0} A_j a^j; \quad A = W_{ab}, S_{ab}, \quad a, b = g, p, \quad (11a)$$

Equation (11a), via the formalism of Paper I [cf. Eqs. (I.9) and (I.10)], yields the corresponding expansions for the eigenvalues and eigenvectors,

$$B = \sum_{j=0} B_j a^j; \quad B = E_a, Z_{ab}; \quad a, b = g, p. \quad (11b)$$

Equations (10) provide a logical point of departure for the formal development of the PV expansion. In our numerical procedure, however, a simpler route will be followed which does not require such a detailed knowledge of the equilibrium structure of the fluid. We will return to this point in Section V.

#### IV. Determination of $E_j$ and $Z_j$

##### A. The Zero-order Solutions

Equation (I.18a) for  $Z_0$  and  $E_0$  is

$$W_0 Z_0 = S_0 Z_0 E_0 \quad (12)$$

where all matrices pertain to the adiabatic ( $a=0$ ,  $p=p_0$ ,  $\varrho=\varrho_0$ ) reference fluid. From Equations (III.20), (III.15)–(III.17) and/or Equations (A6) and (A7), one observes the

following form of the  $W_0$  and  $S_0$  matrices:

$$W_0 = \begin{bmatrix} 0 & 0 \\ 0 & W_{0pp} \end{bmatrix}, \quad (13a)$$

$$S_0 = \begin{bmatrix} S_{0gg} & 0 \\ 0 & S_{0pp} \end{bmatrix}, \quad (13b)$$

where  $W_{0pp}$ ,  $S_{0gg}$ ,  $S_{0pp}$  are real symmetric, and  $S_{0gg}$ ,  $S_{0pp}$  are in addition positive definite. Henceforth, most matrices will carry a three-character subscript: The first character, an arabic numeral, will indicate the perturbation order; the second and third characters, the letters  $g$  and  $p$ , will indicate the block-specifications of the matrix. Substitute Equations (13) in Equation (12) and block-partition the resulting expression:

$$\begin{bmatrix} 0 & 0 \\ W_{0pp} Z_{0pg} & W_{0pp} Z_{0pp} \end{bmatrix} = \begin{bmatrix} S_{0gg} Z_{0gg} E_{0g} & S_{0gg} Z_{0gp} E_{0p} \\ S_{0pp} Z_{0pg} E_{0g} & S_{0pp} Z_{0pp} E_{0p} \end{bmatrix}. \quad (14)$$

From the  $gg$ -,  $gp$ - and  $pg$ -blocks of Equation (14), one obtains

$$E_{0g} = 0, \quad (15)$$

$$Z_{0gp} = 0, \quad (16a)$$

$$Z_{0pg} = 0, \quad (16b)$$

respectively. The  $pp$ -block of Equation (14) gives

$$W_{0pp} Z_{0pp} = S_{0pp} Z_{0pp} E_{0p}. \quad (17a)$$

Equation (17a) is an eigenvalue equation for  $Z_{0pp}$  and  $E_{0p}$ . The eigenvalues are the roots of the secular determinant [cf. Eq. (I.18c)],

$$|W_{0pp} - \varepsilon_{0p} S_{0pp}| = 0, \quad (17b)$$

and the orthonormality condition is [cf. Eq. (I.18b)]

$$Z_{0pp}^\dagger S_{0pp} Z_{0pp} = I. \quad (17c)$$

Equations (17) give complete information on the zeroth-order  $p$ -modes. For the  $g$ -modes,  $E_{0g}$  vanishes because of vanishing  $W_{0gg}$ ; the  $Z_{0gg}$  remains undetermined at this stage. We may summarize the results for  $E_0$  and  $Z_0$  as follows:

$$E_0 = \begin{bmatrix} 0 & 0 \\ 0 & E_{0p} \end{bmatrix}; \quad E_{0p} \text{ diagonal}, \quad (18a)$$

$$Z_0 = \begin{bmatrix} Z_{0gg} & 0 \\ 0 & Z_{0pp} \end{bmatrix}; \quad Z_{0gg} \text{ undetermined}. \quad (18b)$$

##### B. The First-order Solution

Equations (I.19) for  $Z_1$  and  $E_1$  are

$$W_0 Z_1 - S_0 Z_1 E_0 - S_0 Z_0 E_1 = -W_1 Z_0 + S_1 Z_0 E_0, \quad (19a)$$

$$Z_0^\dagger S_0 Z_1 + Z_1^\dagger S_0 Z_0 = -Z_0^\dagger S_1 Z_0, \quad (19b)$$

where, in the present problem,  $W_1$  and  $S_1$  are real symmetric matrices; this property, however, is not re-

quired in our formalism. Substitute Equations (13) and (18) in Equations (19) and partition the resulting expressions. Thus,

$$\left[ \begin{array}{c|c} -S_{0gg}Z_{0gg}E_{1g} & -S_{0gg}Z_{1gp}E_{0p} \\ \hline W_{0pp}Z_{1pg} & W_{0pp}Z_{1pp} - S_{0pp}Z_{1pp}E_{0p} - S_{0pp}Z_{0pp}E_{1p} \end{array} \right] = \left[ \begin{array}{c|c} -W_{1gg}Z_{0gg} & -W_{1gp}Z_{0pp} + S_{1gp}Z_{0pp}E_{0p} \\ \hline -W_{1pg}Z_{0gg} & -W_{1pp}Z_{0pp} + S_{1pp}Z_{0pp}E_{0p} \end{array} \right], \quad (20a)$$

and

$$\left[ \begin{array}{c|c} Z_{0gg}^\dagger S_{0gg} Z_{1gg} + Z_{1gg}^\dagger S_{0gg} Z_{0gg} & Z_{0gg}^\dagger S_{0gg} Z_{1gp} + Z_{1gp}^\dagger S_{0pp} Z_{0pp} \\ \hline Z_{0pp}^\dagger S_{0pp} Z_{1pg} + Z_{1pg}^\dagger S_{0gg} Z_{0gg} & Z_{0pp}^\dagger S_{0pp} Z_{1pp} + Z_{1pp}^\dagger S_{0pp} Z_{0pp} \end{array} \right] = - \left[ \begin{array}{c|c} Z_{0gg}^\dagger S_{1gg} Z_{0gg} & Z_{0gg}^\dagger S_{1gp} Z_{0pp} \\ \hline Z_{0pp}^\dagger S_{1pg} Z_{0gg} & Z_{0pp}^\dagger S_{1pp} Z_{0pp} \end{array} \right]; \quad (20b)$$

from Appendix B of Paper I, we recall that only the diagonal elements of Equation (20b) contain independent information to supplement Equation (20a). Let us discuss Equations (20) block-wise:

i) The  $gg$ -block of Equation (20a) gives

$$W_{1gg}Z_{0gg} = S_{0gg}Z_{0gg}E_{1g}. \quad (21a)$$

Equation (21a) is an eigenvalue equation for  $Z_{0gg}$  with  $E_{1g}$  as its corresponding eigenvalues. The latter are the root of the secular equation

$$|W_{1gg} - \varepsilon_{1g}S_{0gg}| = 0. \quad (21b)$$

The orthonormality condition is [cf. Eq. (I.18b)]

$$Z_{0gg}^\dagger S_{0gg} Z_{0gg} = I. \quad (21c)$$

Thus, the  $gg$ -block of Equation (20a) specifies  $Z_{0gg}$  which was undetermined in the zeroth order; on the other hand, as we shall see,  $Z_{1gg}$  will remain unspecified until the second-order calculations.

ii) The  $gp$ -block of Equation (20a) is

$$S_{0gg}Z_{1gp}E_{0p} = W_{1gp}Z_{0pp} - S_{1gp}Z_{0pp}E_{0p}, \quad (22a)$$

which yields

$$Z_{1gp} = S_{0gg}^{-1} [W_{1gp}Z_{0pp}E_{0p}^{-1} - S_{1gp}Z_{0pp}]. \quad (22b)$$

iii) The  $pg$ -block of Equation (20a) is

$$W_{0pp}Z_{1pg} = -W_{1pg}Z_{0gg}, \quad (23a)$$

from which, via Equation (17a), one gets

$$Z_{1pg} = -Z_{0pp}E_{0p}^{-1}Z_{0pp}^\dagger W_{1pg}Z_{0gg}. \quad (23b)$$

iv) The  $pp$ -block of Equation (20a) is

$$\begin{aligned} W_{0pp}Z_{1pp} - S_{0pp}Z_{1pp}E_{0p} - S_{0pp}Z_{0pp}E_{1p} \\ = -W_{1pp}Z_{0pp} + S_{1pp}Z_{0pp}E_{0p}. \end{aligned} \quad (24a)$$

Equation (24a) must be supplemented with the diagonal elements of the  $pp$ -block of the orthonormality Equation (20b):

$$[Z_{0pp}^\dagger S_{0pp} Z_{1pp}]^{\text{diagonal}} = -\frac{1}{2} [Z_{0pp}^\dagger S_{1pp} Z_{0pp}]^{\text{diagonal}}. \quad (24b)$$

Equations (24) for  $E_{1p}$  and  $Z_{1pp}$  are identical with Equations (I.19) and (I.22). Therefore, from Equations

(I.21), (I.24b), (I.23b) and (I.23a), one obtains, respectively:

$$E_{1p} = [Z_{0pp}^\dagger (W_{1pp}Z_{0pp} - S_{1pp}Z_{0pp}E_{0p})]^{\text{diagonal}}, \quad (25a)$$

$$Z_{1pp} = Z_{0pp}Q_{1pp}, \quad (25b)$$

where

$$Q_{1pp}^{\text{diagonal}} = -\frac{1}{2} [Z_{0pp}^\dagger S_{1pp} Z_{0pp}]^{\text{diagonal}}, \quad (25c)$$

$$\begin{aligned} Q_{1pp}^{rs} = (\varepsilon_{0p}^s - \varepsilon_{0p}^r)^{-1} [Z_{0pp}^\dagger W_{1pp} Z_{0pp} \\ - Z_{0pp}^\dagger S_{1pp} Z_{0pp} E_{0p}]^{rs}; \quad r \neq s. \end{aligned} \quad (25d)$$

Equations (21)–(25) summarize all the information that one may extract from the first-order Equations (20); as previously anticipated,  $Z_{1gg}$  is not determined in this order, and accordingly, the  $gg$ -block of the orthogonality Equation (20b) is not yet employed. The pattern of emergence of solutions is the following: From the  $j^{\text{th}}$ -order perturbation equation,  $E_{jg}$ ,  $E_{jp}$ ,  $Z_{jgp}$ ,  $Z_{jpg}$  and  $Z_{jpp}$  are determined;  $Z_{jgg}$ , however, must be determined from the  $gg$ -block of the perturbation equation in the  $(j+1)^{\text{th}}$  order and the  $gg$ -block of the orthonormality relation in the  $j^{\text{th}}$  order. This is a consequence of the vanishing of  $W_{0gg}$ ; therefore, to determine  $E_{2g}$  and  $Z_{1gg}$ , the second-order perturbation equations are explored.

### C. The Second-order Solutions

From Equations (I.17a), one has

$$\begin{aligned} W_0Z_2 + W_1Z_1 - S_0Z_2E_0 - S_0Z_0E_2 - S_1Z_1E_0 - S_0Z_1E_1 \\ = -W_2Z_0 + S_2Z_0E_0 + S_1Z_0E_1; \end{aligned} \quad (26)$$

only  $E_{2g}$  and  $Z_{1gg}$  will be determined from this equation. The  $gg$ -block of Equation (26), after some rearrangement of terms, yields

$$\begin{aligned} W_{1gg}Z_{1gg} - S_{0gg}Z_{1gg}E_{1g} - S_{0gg}Z_{0gg}E_{2g} \\ = -W_{2gg}Z_{0gg} - W_{1gp}Z_{1pg} + S_{1gg}Z_{0gg}E_{1g}. \end{aligned} \quad (27a)$$

Equation (27a) must be supplemented with the as yet unused diagonal  $gg$ -elements of Equation (20b), the orthonormality condition in the first order:

$$[Z_{0gg}^\dagger S_{0gg} Z_{1gg}]^{\text{diagonal}} = -\frac{1}{2} [Z_{0gg}^\dagger S_{1gg} Z_{0gg}]^{\text{diagonal}}. \quad (27b)$$

Equations (27), except for the additional known interaction term,  $W_{1gp}Z_{1pg}$ , on the right side of Equation

(27a), are the same as Equations (I.17). Following the procedure of Equations (I.26)–(I.29), one gets

$$E_{2g} = [Z_{0gg}^\dagger (W_{2gg} Z_{0gg} + W_{1gp} Z_{1pg}) - S_{1gg} Z_{0gg} E_{1g}]^{\text{diagonal}}, \quad (28a)$$

$$Z_{1gg} = Z_{0gg} Q_{1gg}, \quad (28b)$$

where

$$Q_{1gg}^{\text{diagonal}} = -\frac{1}{2} [Z_{0gg}^\dagger S_{1gg} Z_{0gg}]^{\text{diagonal}}, \quad (28c)$$

$$Q_{1gg}^{rs} = (\varepsilon_{1g}^s - \varepsilon_{1g}^r)^{-1} [Z_{0gg}^\dagger (W_{2gg} Z_{0gg} + W_{1gp} Z_{1pg}) - S_{1gg} Z_{0gg} E_{1g}]^{rs}; \quad r \neq s. \quad (28d)$$

We conclude this Section by a recapitulation of the salient points:

a) The  $p$ -modes in the zeroth order (i.e.,  $Z_{0pp}$  and  $E_{0p}$ ) are determined by the eigenvalue Equations (17).

b) For the  $g$ -modes in this order,  $E_{0g}$  is zero;  $Z_{0gg}$  is obtained from the eigenvalue Equations (21); the matrix entering the latter equation is  $W_{1gg}$ , rather than  $W_{0gg}$ , which has vanished identically. Therefore, the eigenvalues corresponding to  $Z_{0gg}$  are  $E_{1g}$ ; these first-order quantities are the growth rate of the  $g$ -eigenvalues with the non-adiabaticity,  $a$ ; also, in view of vanishing  $E_{0g}$ ,  $aE_{1g}$  is the first non-vanishing term in the  $g$ -mode perturbation expansion. Equations (21) are among the main findings of this paper and, to the authors' knowledge, have not been previously deduced.

c) The  $g$ - and  $p$ -modes are analytically separated in the zeroth order, i.e.,  $Z_{0gp}$  and  $Z_{0pg}$  vanish. This is a consequence of Sobouti's definition of the  $g$ - and  $p$ -modes (cf. the Propositions 1 and 2 of Paper III); the definition leads to the two independent basis sets  $\{\zeta_g^u\}$  and  $\{\zeta_p^v\}$  of which the former spans the  $g$ -subspace and the latter spans the  $p$ -subspace of the normal modes of the adiabatic fluid.

d) The  $p$ -modes in the first order are determined from Equations (25a), (22b) and (25b).

e) For the  $g$ -modes in this order,  $Z_{1gg}$  and  $Z_{1pg}$  are solutions of Equations (28b) and (23b), respectively; the interaction between the two modes appears in the first order. The second-order  $g$ -eigenvalue,  $E_{2g}$ , is obtained from Equation (28a). We stress that the normal-mode analysis of self-gravitating fluids has been reduced 1) to solution of two ordinary uncoupled eigenvalue-Equations (17) and (21) in the zeroth order; and 2) to solution of a set of linear inhomogeneous algebraic equations of the type of Equations (23)–(25) in the higher orders.

Finally, let us recall that the energy denominator, which always appears in the calculation of the first- and higher-order perturbation expansions of eigenvectors, is embedded in the  $Q$ -matrices of Equations (25d) and (28d) and is also present in  $Z_{1gp}$  and  $Z_{1pg}$  of Equations (22b) and (23b), respectively; indeed, the latter two equations can be written in terms of  $Q$ -matrices which are simpler in form due to the vanishing of  $E_{0g}$ .

## V. Numerical Procedure

Recall that in Section III, we regarded the expansions of  $g$ ,  $p$ , and their derivatives in powers of the non-adiabaticity parameter  $a$ , Equations (10), as the logical starting point for the subsequent development of the PV expansions. Despite this possibility, we note that we have actually only required the perturbation expansions of the  $W$ - and  $S$ -matrices in the preceding analysis. Thus in our numerical procedure, we have directly obtained the expansions of the matrices from Equations (A6) and (A7) of Appendix A: The  $W$ - and  $S$ -elements given in these equations were calculated for seven polytropes of indices  $n_0 = 1.5$  and  $n_{\pm i} = n_0 \pm i\Delta n$ ,  $i = 1, 2, 3$  (note that the polytrope 1.5 is the adiabatic fluid corresponding to the ratio of specific heats 5/3). The resulting expressions were then substituted in the following finite-difference equations to obtain the desired expansions:

$$\left(\frac{df}{dn}\right)_{n_0} = [(f_3 - f_{-3}) - 9(f_2 - f_{-2}) + 45(f_1 - f_{-1})]/60\Delta n + \mathcal{O}[(\Delta n)^7], \quad (29a)$$

$$\left(\frac{d^2f}{dn^2}\right)_{n_0} = [(f_3 + f_{-3}) - \frac{27}{2}(f_2 + f_{-2}) + 135(f_1 + f_{-1}) - 245f_0]/90\Delta n + \mathcal{O}[(\Delta n)^8], \quad (29b)$$

where  $f$  is any element of the  $W$ - and  $S$ -matrices, and the indices in  $f$  indicate the polytropic index at which  $f$  is evaluated. The derivatives with respect to  $n$  were then transformed to those with respect to  $a$  by the defining Equation (9c) and the Taylor expansions of  $W$  and  $S$ , Equations (11a), were constructed.

The polytropic data were generated directly. For  $\xi \leq 0.3\xi_0$ , where  $\xi$  and  $\xi_0$  are the polytropic radius variable and its surface value, respectively, a Taylor expansion about the origin was employed; the expansion contained six terms. For  $0.3\xi_0 < \xi \leq \xi_0$ , an extrapolation formula, containing seven terms, was used. The second and higher derivatives of the polytropic temperature were obtained from the Lane-Emden equation and from successive differentiation of the latter. The numerical results were compared with the values of the British Association Tables (1932). In the overlapping region, the two results agreed up to eight figures, the accuracy of the aforementioned Tables. After the polytropic temperature, density, and pressure were obtained, the matrix elements were integrated by Simpson's rule.

## VI. Conclusions and Numerical Results

Strictly speaking, eigenvalues and eigenvectors are solutions of secular equations, e.g.,  $E_{0p}$ ,  $Z_{0pp}$  and  $E_{1g}$ ,  $Z_{0gg}$ , of Equations (17) and (21), respectively. As a matter of convenience, however, in this Section we shall loosely refer to any  $j^{\text{th}}$ -order expansion term of these quantities as the  $j^{\text{th}}$ -order eigenvalue and eigenvector.





of the  $W_{1gg}$ -matrix is positive definite; therefore all the corresponding eigenvalues,  $E_{1g}$ , are negative. As the mode number increases, the increasing sequence of the negative eigenvalues accumulates at zero; thus, the spectrum of  $E_{1g}$  closely resembles the energy spectrum of the hydrogen atom. An inspection of Tables 1 and 2 in various RR-approximations reveals that the  $E_{1g}$  eigenvalues obey the minimum principle; as the approximation order increases, the eigenvalues *approach their exact limits from above* and strictly obey the interleaving rule of Equation (I.8). Note the rapidity of the variational convergence of the lower-order  $g$ -modes.

The zeroth-order  $p$ -eigenvectors are solutions of the eigenvalue Equation (17a);  $W_{0pp}$ , the matrix entering this equation, is positive definite; therefore all the eigenvalues,  $E_{0p}$ , are positive. As the mode order increases, the increasing sequence of the positive eigenvalues accumulates at infinity; the spectrum of  $E_{0p}$  is thus like the energy spectrum of a harmonic oscillator. Again one can see from Tables 1 and 2 that  $E_{0p}$  obeys the minimum principle; as the RR-approximation order increases, the  $p$ -eigenvalues *approach their limits from above* and satisfy the interleaving rule of Equation (I.8). Also, note the rapidity of variational convergence of the lower-order  $p$ -modes.

In a non-adiabatic fluid, where an analytical separation of the two modes is not available, such clear-cut behavior of the eigenvalues, that is, whether or not they approach the limits monotonically from above, is not observed: In the adiabatic case, the unambiguous convergence of one mode-type can be examined independently of the other; in the non-adiabatic case, however, the provision for examining such independent convergence is lost due to the loss of the orthogonality of the two trial sets,  $\{\zeta_g\}$  and  $\{\zeta_p\}$ . Nevertheless, even in the non-adiabatic case, both modes, when combined together, still satisfy the minimum principle in that a) the smallest eigenvalue of a given RR-approximation is smaller than the smallest eigenvalue of the preceding approximation, and b) when arranged in an ascending sequence, the eigenvalues of two successive RR-approximations interleave in compliance with Equation (I.8). To demonstrate this behavior, we have compiled Table 3 from Sobouti's (1977b) data; the Table contains both the  $g$ - and  $p$ -eigenvalues of polytropes 1 and 2 for  $l=2$  in three RR-approximations with 6, 7 and 8 variational parameters. The columns of Table 3 are arranged in such a way that the largest eigenvalue of a given sequence is at the top. Interleaving of the eigenvalues is evident. It is interesting to note, however, that in no case do  $g$ -eigenvalues interleave  $p$ -eigenvalues, and vice-versa. This last feature is attributable to the fact that the trial basis sets,  $\{\zeta_g\}$  and  $\{\zeta_p\}$ , bear a fair resemblance to the exact eigenvectors, even in the non-adiabatic case, and remain a flexible tool for investigating their further properties (see Paper III, Section VIII). These comments should throw some light on the long-standing question of

**Table 3.** A demonstration of the variational principle as obeyed by the  $g$ - and  $p$ -eigenvalues of polytropes 1 and 2 for  $l=2$ : a) Lowest eigenvalue in a given RR-approximation is lower than the lowest eigenvalue of the preceding approximation; and b) the eigenvalues in two successive approximations interleave. Data are taken from Sobouti (1977b)

| RR-approximation order           |            |            |            |
|----------------------------------|------------|------------|------------|
|                                  | (6)        | (7)        | (8)        |
| Superadiabatic fluid: $n=1, l=2$ |            |            |            |
| $p_4$                            |            |            | 9.402720   |
| $p_3$                            | 5.053224   | 5.053257   | 4.773653   |
| $p_2$                            | 1.888032   | 1.888047   | 1.885822   |
| $p_1$                            | 0.3035501  | 0.3035503  | 0.3035496  |
| $g_4$                            |            | -0.0097782 | -0.0100069 |
| $g_3$                            | -0.0158273 | -0.0161557 | -0.0161594 |
| $g_2$                            | -0.0278728 | -0.0279221 | -0.0280226 |
| $g_1$                            | -0.0613717 | -0.0613774 | -0.0613814 |
| Subadiabatic fluid: $n=2, l=2$   |            |            |            |
| $p_4$                            |            |            | 4.906920   |
| $p_3$                            | 2.750171   | 2.750247   | 2.183195   |
| $p_2$                            | 1.027453   | 1.027481   | 1.013811   |
| $p_1$                            | 0.2730481  | 0.2730486  | 0.2730481  |
| $g_1$                            | 0.0495142  | 0.0495145  | 0.0494278  |
| $g_2$                            | 0.0259738  | 0.0261176  | 0.0260695  |
| $g_3$                            | 0.0139640  | 0.0154700  | 0.0153509  |
| $g_4$                            |            | 0.0078844  | 0.0078387  |

the variational behavior of the  $g$ - and  $p$ -modes; see, for example, Ledoux (1974) and Sobouti (1977a).

In connection with convection in stellar interiors,  $E_{1g}$  and  $Z_{0gg}$  are of particular importance; from the former, one deduces the time rate of growth of the convective instabilities, while from the latter, via Equation (6a), one



obtains the patterns of the convective motions. With regard to the stability of the fluid against small departures from equilibrium, following Sobouti (Paper III, Section IX), let us argue that the fluid will be stable if and only if all of its normal modes are stable: The  $p$ -modes are always stable; the  $g$ -modes will be stable if the eigenvalues,  $aE_{1g}$ , are all positive. This condition, in view of the negative values of  $E_{1g}$ , reduces to  $a$  being negative, that is, to the fluid being superadiabatic; this is, of course, Schwarzschild's criterion for convective stability. For further discussion of this issue, see Equation (III.26) and accompanying remarks.

A comment on the notation is appropriate: Because of the vanishing of the zeroth-order  $g$ -eigenvalues,  $E_{1g}$  and  $E_{2g}$  bear the same relation to  $Z_{0gg}$  and  $Z_{1gg}$ , respectively, as  $E_{0p}$  and  $E_{1p}$  to their respective eigenvectors; this may suggest relabelling  $E_{1g}$  and  $E_{2g}$  as  $E_{0g}$  and  $E_{1g}$ , respectively, to indicate their parental relation to the corresponding eigenvectors. We have refrained from doing this, however, because it would divest the whole PV formalism of its symmetry by setting up two different ways of ordering the perturbation terms for the  $g$ - and  $p$ -eigenvalues; this is reflected physically by the necessity of forming  $aE_{1g}$  and  $a^2E_{2g}$  to evaluate the perturbation expansions, Equation (11b).

We conclude this Section by noting that many features discussed above, e.g., separation of the two types of modes, their variational behavior in the zeroth order, and their subsequent interaction in the higher perturbation orders, would not have been revealed without a PV expansion procedure such as we have adopted; the facility of computation, developed formally in Paper I, and extended and utilized in the present paper, is worthy of attention.

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## Appendix A: The Matrix Elements of Equations (6) and (7)

The  $W$ -matrix. After some integrations by parts and the use of Equation (2b), Equations (8) and (1b) reduce to

$$W_{ab}^{ts} = - \int \nabla \cdot \zeta_a^t \delta_b^s p dv - \int \frac{1}{2} \zeta_a^t \cdot \nabla p \delta_b^s \varrho dv - \int \delta_a^t \varrho \delta_b^s \Omega dv; \quad a, b = g, p. \quad (A1)$$

The  $p$ -basis vectors of Equation (5b) are the same as those of Paper III; therefore, the  $pp$ -elements of  $W$ - and  $S$ -matrices will remain as in the latter paper. The  $g$ -basis vectors, however, are somewhat different and result in different expressions for  $W_{ga}$  and  $S_{ga}$ ,  $a = g, p$ . Substitute

$\zeta_g^s$  of Equation (5a) in Equations (2a) and (2b) to obtain

$$\delta_g^s p = \gamma p \nabla \cdot \zeta_g^s \left[ \frac{d \ln p / dr}{d \ln p_0 / dr} - 1 \right], \quad (A2)$$

$$\delta_g^s \varrho = \varrho \nabla \cdot \zeta_g^s \left[ \frac{d \ln \varrho / dr}{d \ln \varrho_0 / dr} - 1 \right]. \quad (A3)$$

We note that as the fluid tends to the adiabatic limit (i.e., as  $p, \varrho \rightarrow p_0, \varrho_0$ ),  $\delta_g^s p$  and  $\delta_g^s \varrho$ , and consequently  $W_{gg}$ ,  $W_{gp}$  and  $W_{pg}$  all tend to zero.

Let the basis vectors  $\{\zeta_g^s\}$  and  $\{\zeta_p^s\}$  have the following spherical harmonic expansions [cf. Eqs. (III.5) and (III.7)]:

$$\zeta_a^s = \left( \frac{1}{r^2} \psi_a^s Y_l^m, \frac{1}{l(l+1)} \frac{\chi_a^s}{r} \frac{\partial Y_l^m}{\partial \theta}, \frac{1}{l(l+1)} \frac{\chi_a^s}{r} \frac{1}{\sin \theta} \frac{\partial Y_l^m}{\partial \phi} \right); \quad (A4)$$

$$a = g, p.$$

The constraints of Equations (5) enable one to eliminate one of the two scalars  $\psi_a^s$  or  $\chi_a^s$  in favor of the other (see Paper III). Thus,

$$\chi_g^s = \psi_g^s + \frac{\varrho'_0}{\varrho_0} \psi_g^s, \quad (A4a)$$

$$\chi_p^s = l(l+1) \frac{\psi_p^s}{r^2}. \quad (A4b)$$

Also note that

$$\nabla \cdot \zeta_g^s = - \frac{\varrho'_0}{\varrho_0} \frac{\psi_g^s}{r^2} Y_l^m, \quad (A5a)$$

$$\nabla \cdot \zeta_p^s = \frac{1}{r^2} (\psi_p^s - \chi_p^s) Y_l^m. \quad (A5b)$$

Substitution of Equations (A2)–(A5) in Equation (A1) and reductions parallel to those of Paper III lead to the following matrix elements:

$$W_{gg}^{ts} = W_{gg}^{st} = - \int_0^R \frac{\varrho'_0}{\varrho_0} p \left[ \frac{p'}{p} - \frac{p'_0}{p_0} \right] \psi_g^t \psi_g^s \frac{dr}{r^2} + \int_0^R p' \left[ \frac{\varrho'}{\varrho} - \frac{\varrho'_0}{\varrho_0} \right] \psi_g^t \psi_g^s \frac{dr}{r^2} - 4\pi G \int_0^R Y_g^t Y_g^s dr, \quad (A6a)$$

$$W_{gp}^{ts} = W_{pg}^{st} = - \int_0^R p \left[ \frac{p'}{p} - \frac{p'_0}{p_0} \right] \psi_g^t (\psi_p^s - \chi_p^s) \frac{dr}{r^2} + \int_0^R p' \left[ \frac{\varrho'}{\varrho} - \frac{\varrho'_0}{\varrho_0} \right] \psi_g^t \psi_p^s \frac{dr}{r^2} - 4\pi G \int_0^R Y_g^t Y_p^s dr, \quad (A6b)$$

and from Equations (III.15c),

$$W_{pp}^{ts} = W_{pp}^{st} = \int_0^R \gamma p (\psi_p^t - \chi_p^t) (\psi_p^s - \chi_p^s) \frac{dr}{r^2} + \int_0^R p' \left[ \psi_p^t (\psi_p^s - \chi_p^s) + \psi_p^s (\psi_p^t - \chi_p^t) - \frac{\varrho'}{\varrho} \psi_p^t \psi_p^s \right] \frac{dr}{r^2} - 4\pi G \int_0^R Y_p^t Y_p^s dr, \quad (A6c)$$

where

$$Y_g^t = -r^l \int_r^R \varrho \left[ \frac{\varrho'}{\varrho} - \frac{\varrho'_0}{\varrho_0} \right] \psi_g^t \frac{dr}{r^{l+1}}, \quad (\text{A6d})$$

and from Equations (III.16b) and (III.16c),

$$Y_p^t = \varrho \frac{\psi_p^t}{r} - r^l \int_r^R \varrho \left[ (l+1) \frac{\psi_p^t}{r} - \chi_p^t \right] \frac{dr}{r^{l+1}}. \quad (\text{A6e})$$

The  $S$ -matrix. From Equations (7b) and (A4), one obtains

$$S_{gg}^{ts} = S_{gg}^{st} = \int_0^R \varrho \left[ \frac{1}{r^2} \psi_g^t \psi_g^s + \frac{1}{l(l+1)} \chi_g^t \chi_g^s \right] dr, \quad (\text{A7a})$$

$$S_{gp}^{ts} = S_{gp}^{st} = - \frac{1}{l(l+1)} \int_0^R \varrho \left[ \frac{\varrho'}{\varrho} - \frac{\varrho'_0}{\varrho_0} \right] \psi_g^t \chi_p^s dr, \quad (\text{A7b})$$

$$S_{pp}^{ts} = S_{pp}^{st} = \int_0^R \varrho \left[ \frac{1}{r^2} \psi_p^t \psi_p^s + \frac{1}{l(l+1)} \chi_p^t \chi_p^s \right] dr, \quad (\text{A7c})$$

where Equations (A7a) and (A7c) are the same as Equations (III.17a) and (III.17c). This completes determination of the  $W$ - and  $S$ -matrices.

## Appendix B: Special Provision for the Solid-body Translation of the Fluid

From Section VII of Paper III, we recall that the lowest  $p$ -mode for  $l=1$  is a neutral Kelvin mode and represents a solid-body translation of the fluid; this mode, denoted by a superscript  $I$ , has the exact solution,

$$\psi_p^I = \chi_p^I = r^2; \quad l=1. \quad (\text{B1})$$

This vector has been chosen to be the first member of the  $p$ -basis vectors of Equations (A4b), and being a neutral mode, renders the first column and first row of the  $W_{0pp}$ -matrix zero. Thus,  $W_{0pp}$ , and consequently  $E_{0p}$ , become singular, and Equations (22b) and (23b) will no longer be adequate to calculate  $Z_{1gp}$  and  $Z_{1pg}$ . To obtain these quantities in this special case, the following direct analysis of Equations (22a) and (23a) is carried out: Where applicable, separate the *first  $p$ -column* and/or the *first  $p$ -row* of any matrix, and in the notation of Equation (B1), denote them by a superscript  $I$ ; denote the *remainder of the matrix* in question by a superscript  $R$ . One obtains

$$W_{0pp} = \begin{bmatrix} 0 & 0 \\ 0 & W_{0pp}^{RR} \end{bmatrix}, \quad (\text{B2a})$$

$$W_{1gp} = (W_{1pg})^\dagger = [0 \mid W_{1gp}^R], \quad (\text{B2b})$$

$$S_{0pp} = \begin{bmatrix} S_{0pp}^{II} & S_{0pp}^{IR} \\ S_{0pp}^{RI} & S_{0pp}^{RR} \end{bmatrix}, \quad (\text{B2c})$$

$$S_{1pg} = (S_{1gp})^\dagger = \begin{bmatrix} S_{1pg}^I \\ S_{1pg}^R \end{bmatrix}, \quad (\text{B2d})$$

where  $W_{0pp}^{RR}$  is the matrix block generated from the  $p$ -basis vectors other than the first  $p$ -vector, and  $W_{1gp}^R$  is the

matrix block generated by a  $g$ -vector and a  $p$ -vector other than the first  $p$ -vector. The blocks of the  $S$ -matrix have a similar interpretation. Correspondingly, the  $E$  and  $Z$  matrices assume the following form:

$$E_{0p} = \begin{bmatrix} 0 & 0 \\ 0 & E_{0p}^R \end{bmatrix}, \quad E_{0p}^R = \text{diagonal}, \quad (\text{B3a})$$

$$Z_{0pp} = \begin{bmatrix} Z_{0pp}^{II} & Z_{0pp}^{IR} \\ 0 & Z_{0pp}^{RR} \end{bmatrix}, \quad (\text{B3b})$$

$$Z_{1gp} = [0 \mid Z_{1gp}^R], \quad (\text{B3c})$$

$$Z_{1pg} = \begin{bmatrix} Z_{1pg}^I \\ Z_{1pg}^R \end{bmatrix}. \quad (\text{B3d})$$

In Equations (B3b) and (B3c),  $Z_{0pp}^{RI}$  and  $Z_{1gp}^I$  have, respectively, vanished because the solid-body motion of Equation (B1) is an exact solution of the problem and is thus orthogonal to the other modes.

a) *Determination of  $Z_{1gp}$* : Substitute Equations (B2) and (B3) in Equation (22a) and partition the result; one gets

$$S_{0gg} Z_{1gp}^R E_{0p}^R = W_{1gp}^R Z_{0pp}^{RR} - (S_{1gp}^I Z_{0pp}^{IR} + S_{1gp}^R Z_{0pp}^{RR}) E_{0p}^R. \quad (\text{B4})$$

Since both  $S_{0gg}$  and  $E_{0p}^R$  are non-singular, one readily obtains

$$Z_{1gp}^R = (S_{0gg})^{-1} [W_{1gp}^R Z_{0pp}^{RR} (E_{0p}^R)^{-1} - S_{1gp}^I Z_{0pp}^{IR} - S_{1gp}^R Z_{0pp}^{RR}]. \quad (\text{B5})$$

This equation, together with Equation (B3c), completes the solution for  $Z_{1gp}$ .

b) *Determination of  $Z_{1pg}$* : The orthogonality relation governing  $Z_{1pg}$  and  $Z_{1gp}$  provides a simple derivation of the former in terms of the latter. From the  $pg$ -block of Equation (20b), one has

$$Z_{0pp}^\dagger S_{0pp} Z_{1pg} = -Z_{1gp}^\dagger S_{0gg} Z_{0gg} - Z_{0pp}^\dagger S_{1pg} Z_{0gg}. \quad (\text{B6})$$

Reduction of Equation (B6) by Equation (17c) gives

$$Z_{1pg} = -Z_{0pp} (Z_{1gp}^\dagger S_{0gg} + Z_{0pp}^\dagger S_{1pg}) Z_{0gg}. \quad (\text{B7})$$

Alternatively, a direct partitioning of Equation (23a) could lead to  $Z_{1pg}$ , but the above procedure is simpler.

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